# Astrophysics of Compact Objects (171.156), Fall 2011 

## Problem Set 9

Due: 15 November 2011

1. What is the condition on the energies $\epsilon_{1}$ and $\epsilon_{2}$ of two photons that collide at an angle $\theta$ if they are to produce an electron-positron pair?
2. A fireball has a rest mass $M$ (baryons) and energy $E=\eta M$. It begins as a sphere with size $R_{0}$. What is the requirement on $\eta$ so that most of the energy will be converted to kinetic energy of the baryons. (The baryons must be coupled to the photons at least up to the time where the Lorentz factor has saturated.)
3. Write the radius at which the shell will begin to spread (increase its width) as a function of the initial fireball parameters $E, M$, and $R_{0}$.
4. Two shells of similar mass and Lorentz factor $\gamma_{1}$ and $\gamma_{2}$ collide. Use conservation of energy and momentum to find the fraction of energy that will thermalize.
5. Write the expression for the synchrotron self-absorption frequency $\nu_{a}$ as a function of time $t$, fireball energy $E$, ISM particle density $n$, and the fractions $\epsilon_{e}$ and $\epsilon_{B}$ of the fireball energy converted to electron energy and magnetic fields. Separate two cases of fast and slow cooling to determine the effective temperature of the radiating electrons.
6. If the flow is not spherical but occupies a cone with opening angle $\theta$, then the evolution is similar to spherical as long as $\gamma>1 / \theta$. After that the jet spreads with $\theta(t) \sim 1 / \gamma$. When is the deviation from spherical behavior expected if the opening angle of the jet is $\theta=0.1$ ? Use the conservation of energy to show that the radius as function of observed time is almost constant. Find $\nu_{c}, \nu_{m}, \nu_{a}$ and $F_{\max }$ for this case. What will be the typical decay rate of the optical flux?
