

General Relativity (Ph236a)

Problem Set 3

Due: October 17, 2006

Preview: Problem 1 is essential; it defines the Riemann normal coordinates, or the coordinates for a locally inertial frame. Problems 2–5 are exercises intended to develop a facility with vector fields and their dual one-form fields. Problem 6 is a straightforward exercise to develop some facility with metrics; it also illustrates some features of spinning spacetimes that will show up later when we deal with gravitomagnetism and with spinning black holes.

1. **Riemann normal coordinates, or a locally inertial frame:** Show that a coordinate system can always be chosen so that at a given point p , the metric $g_{\mu\nu}$ is canonical and $\partial_\rho g_{\mu\nu} = 0$.

2. (Wald 3.3) **Commutator of vector fields:**

- a. Verify that the commutator,

$$[v, w](f) = v[w(f)] - w[v(f)],$$

where v and w are smooth vector fields, and f is a function, satisfies the linearity and Leibnitz properties, and hence defines a vector field.

- b. Let X , Y , and Z be vector fields on a manifold. Verify that their commutator satisfies the Jacobi identity,

$$[[X, Y], Z] + [[Y, Z], X] + [[Z, X], Y] = 0.$$

- c. Let Y_1, \dots, Y_n be vector fields on an n -dimensional manifold such that at each point p on the manifold, they form a basis of the tangent space V_p . Then, at each point, we may expand each commutator $[Y_\alpha, Y_\beta]$ in this basis, thereby defining the functions $C^\gamma_{\alpha\beta} = -C^\gamma_{\beta\alpha}$ by

$$[Y_\alpha, Y_\beta] = \sum_\gamma C^\gamma_{\alpha\beta} Y_\gamma.$$

Use the Jacobi identity to derive an equation satisfied by $C^\gamma_{\alpha\beta}$.

3. (Wald 3.4) **More on commutators:**

- a. Show that in any coordinate basis, the components of the commutator of two vector fields v and w are given by

$$[v, w]^\mu = \sum_\nu \left(v^\nu \frac{\partial w^\mu}{\partial x^\nu} - w^\nu \frac{\partial v^\mu}{\partial x^\nu} \right).$$

- b. Let Y_1, \dots, Y_n be as in problem 2(c). Let Y^{1*}, \dots, Y^{n*} be the dual basis. Show that the components $(Y^{\gamma*})_\mu$ of $Y^{\gamma*}$ in any coordinate basis satisfy

$$\frac{\partial (Y^{\gamma*})_\mu}{\partial x^\nu} - \frac{\partial (Y^{\gamma*})_\nu}{\partial x^\mu} = \sum_{\alpha, \beta} C^\gamma_{\alpha\beta} (Y^{\alpha*})_\mu (Y^{\beta*})_\nu.$$

[Hint: Contract both sides with $(Y_\sigma)^\mu(Y_\rho)^\nu$. Also, be careful to distinguish between indices that label the vector fields from indices that label their components.]

4. (MTW, Problem 9.6) **Practice with dual bases:** In a three-dimensional space with spherical coordinates r , θ , and ϕ , it is often useful to use, instead of the coordinate basis $\partial/\partial r$, $\partial/\partial\theta$, $\partial/\partial\phi$, the orthonormal basis,

$$e_{\hat{r}} = \frac{\partial}{\partial r}, \quad e_{\hat{\theta}} = \frac{1}{r} \frac{\partial}{\partial \theta}, \quad e_{\hat{\phi}} = \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi}.$$

- a. What is the one-form basis $\{\omega^{\hat{r}}, \omega^{\hat{\theta}}, \omega^{\hat{\phi}}\}$ dual to this tangent-vector basis?
 - b. On the sphere of unit radius, draw pictures of the bases $\{\partial/\partial r, \partial/\partial\theta, \partial/\partial\phi\}$, $\{e_{\hat{r}}, e_{\hat{\theta}}, e_{\hat{\phi}}\}$, $\{dr, d\theta, d\phi\}$ (the one-form basis dual to the coordinate tangent-vector basis), and $\{\omega^{\hat{r}}, \omega^{\hat{\theta}}, \omega^{\hat{\phi}}\}$.
5. (MTW, Problems 9.7-8) **Commutators for Euclidean space in spherical coordinates:**
- a. Evaluate the commutator of the two vector fields $e_{\hat{\theta}} = (1/r)\partial/\partial\theta$ and $e_{\hat{\phi}} = (1/r \sin \theta)\partial/\partial\phi$. Use this to evaluate the components of the tensor C^i_{jk} (of problems 2 and 3) for this two-dimensional vector space.
 - b. In Cartesian coordinates of a 3-dimensional Euclidean space, define three “angular-momentum operators” $L_i = \epsilon_{ijk} x^j \partial_k$. Evaluate the components of the tensor C^i_{jk} for these vector fields.
6. (Wald 2.8b+) **Rotating coordinates:** The line element of special relativity is

$$ds^2 = -dt^2 + dx^2 + dy^2 + dz^2.$$

- a. Find the line element in rotating coordinates, defined by

$$\begin{aligned} t' &= t, \\ x' &= (x^2 + y^2)^{1/2} \cos(\phi - \omega t), \\ y' &= (x^2 + y^2)^{1/2} \sin(\phi - \omega t), \\ z' &= z, \end{aligned}$$

where $\tan \phi = y/x$.

- b. Re-write the line element in terms of cylindrical coordinates r' and ϕ' defined by $x' = r' \cos \phi'$ and $y' = r' \sin \phi'$.